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Introduction

- Auroral emissions arise from energetic particles precipitating into the upper atmosphere.
- Classifying spectra characterises energy input into the ionosphere–thermosphere system.
- Needs subjective expert visual inspection, which does not scale.

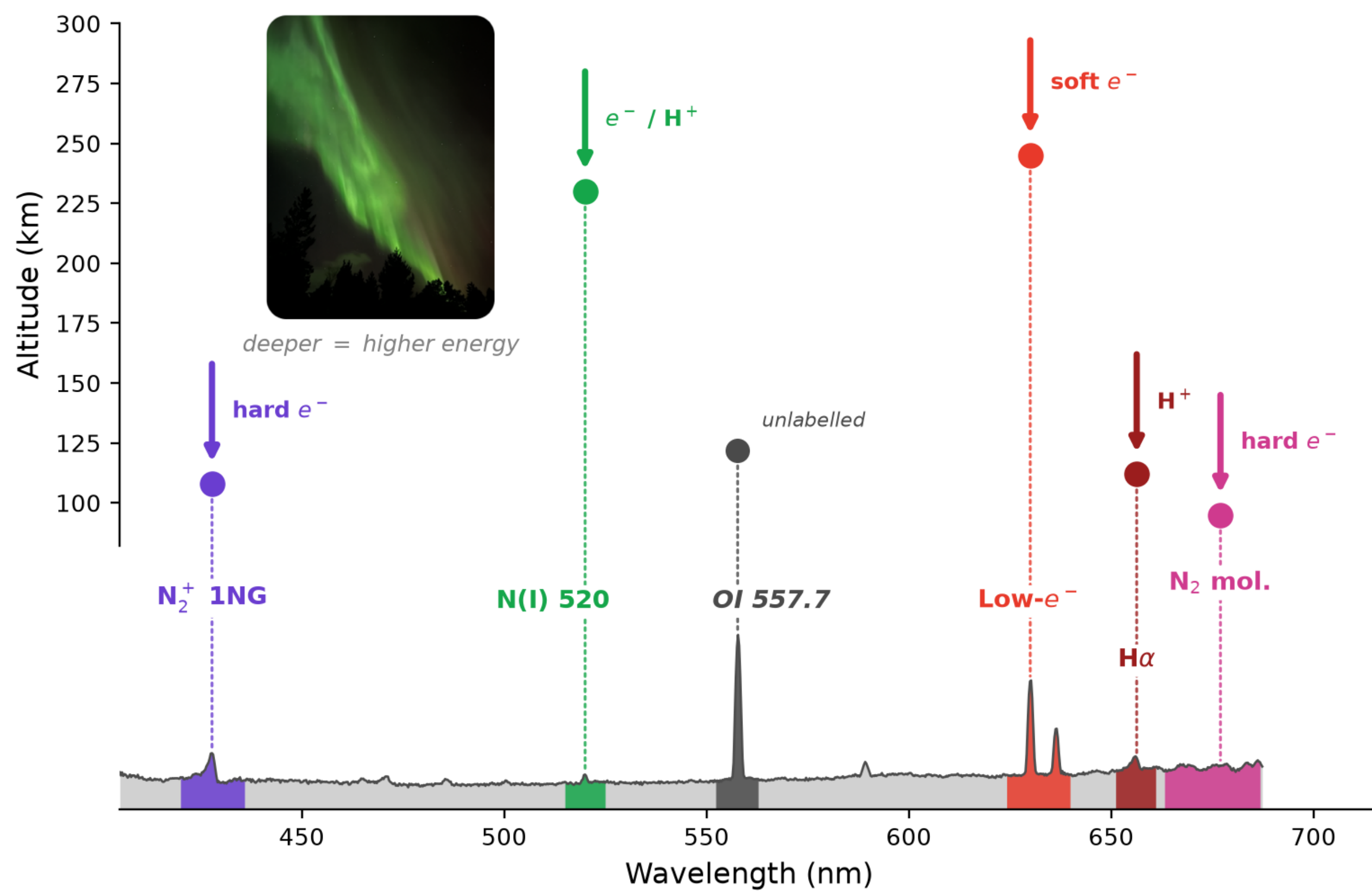


Figure 1. **From precipitation to class label.** Particle energy sets the emission *altitude* and hence the *colour* (top); each line appears in the mean spectrum (bottom) and defines one expert *class*. Soft e^- stay high (red OI 630, Low- e^-); harder e^- go deep (green OI 557.7; violet N_2^+ 427.8); protons give H α . The strong OI 557.7 nm line (dark grey) is present in all spectra and is *not* a label.

Motivation

Can Vision Transformer attention act as a **built-in interpretability** mechanism on ordered 1D scientific signals, where physical ground truth enables direct validation?

Data

- ASIS spectrograph, Skibotn (Norway): fibre-fed Czerny–Turner, 405–687 nm, 1024 bins.
- 719 min-max-normalised spectra, expert-annotated with 5 binary labels.

Table 1. Emission classes, physical tracer and sample count ($N=719$).

Class	Physical tracer	n
H α	Proton precipitation	346
N ₂ mol.	High-energy e^- precipitation	114
Low- e^-	Low-energy e^- precipitation	214
N(I) 520	N(² D → ⁴ S)	307
N ₂ ⁺ 1NG	N ₂ ionisation	267

Method

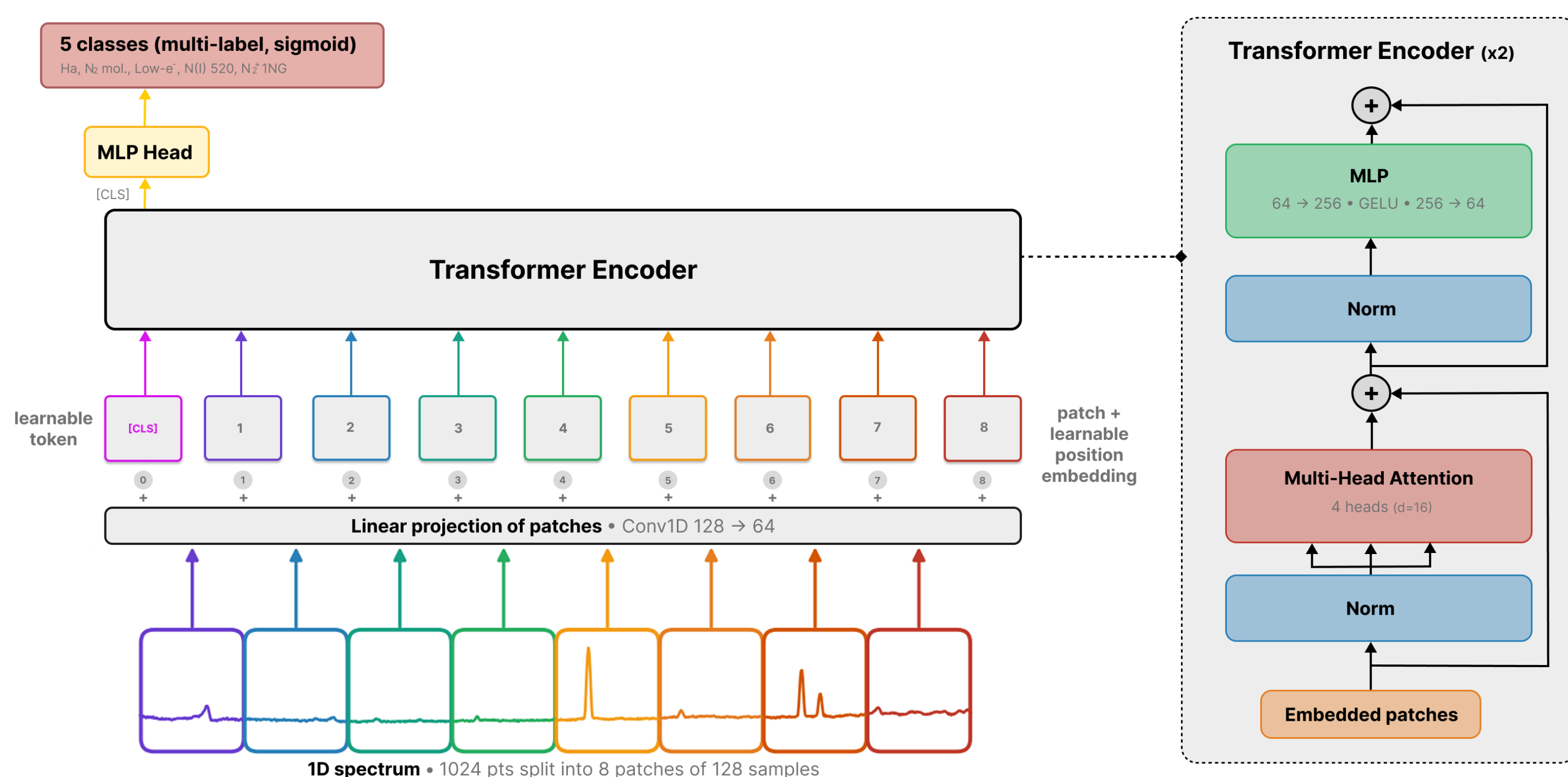


Figure 2. ViT-1D architecture - patch tokenisation, learnable [CLS], Transformer encoder blocks, [CLS] → head.

- We train a **1D Vision Transformer (ViT-1D)** on the raw spectra: patch self-attention with a [CLS] classifier (architecture above).
- We **compare it to an MLP baseline**, the same backbone with a *single full-spectrum patch* (1 layer, 1 head), so attention collapses to a **learned spectral projection**; this isolates the contribution of attention.
- Training & evaluation**: augmentation (noise, scale jitter, channel dropout, Mixup); ViT-1D BCE/AdamW (300 ep.), MLP focal loss (200 ep.); **stratified 5-fold CV** $\times 3$ (15 folds), macro F1 & mAP.

Resources & Demo



Scan for the paper & poster, supplementary material, and a **web app** to run the trained ViT-1D, and inspect its attention maps, on auroral spectra.

Label Co-occurrence

H α	346	0	132	92	4
N ₂ mol.	0	114	7	48	95
Low- e^-	132	7	214	76	9
N(I) 520	92	48	76	307	171
N ₂ ⁺ 1NG	4	95	9	171	267
H α	N ₂ mol.	Low- e^-	N(I) 520	N ₂ ⁺ 1NG	

Figure 3. Label Co-occurrence (data).

Co-occurrence is **set by the precipitating particles**:

- H α & N₂ mol. : 0%, proton vs. high-energy-electron events, *disjoint in this dataset*.
- N(I) 520 & N₂⁺ : 64%. Same particles dissociate and ionise N₂ (co-production).
- N₂ mol. & N₂⁺ : 83%. High-energy electrons make N₂ bands and N₂⁺.

Ground truth

Does the model's **attention** recover this same structure ?

Results

Both architectures reach **strong performance** (macro AP: MLP 81.1%, ViT-1D 77.8%).

Table 2. Per-class F1 and AP (%), mean over 15 folds. Best in **bold**.

Class	F1 (%)		AP (%)	
	MLP	ViT-1D	MLP	ViT-1D
H α	81.8	80.6	84.2	82.7
N ₂ mol.	47.8	52.3	57.4	51.6
Low- e^-	81.1	72.5	86.6	81.9
N(I) 520	75.2	75.4	84.2	80.9
N ₂ ⁺ 1NG	86.0	84.1	93.3	91.9
Macro	74.4	73.0	81.1	77.8

- N₂⁺ 1NG is easiest for both (F1 $\geq$ 84%), a sharp, isolated spectral band.
- N₂ mol. is hardest (F1 $\leq$ 53%): diffuse band structure and only 114 samples. The ViT-1D wins here by **+4.5%** F1, suggesting patch-level attention better captures diffuse band patterns.
- The MLP's largest advantage is on **Low- e^-** (+8.6% F1), whose broad signature benefits from a global receptive field.

Attention as Built-in Interpretability

The [CLS] token's attention is a **saliency map over wavelength**, where the model looks to decide. Per class we contrast positive vs. negative samples:

$$\Delta A_c(\lambda) = \bar{A}_{c=1}(\lambda) - \bar{A}_{c=0}(\lambda),$$

red ($\Delta A_c > 0$) marks where attention *increases* for the class, **blue** where it decreases.

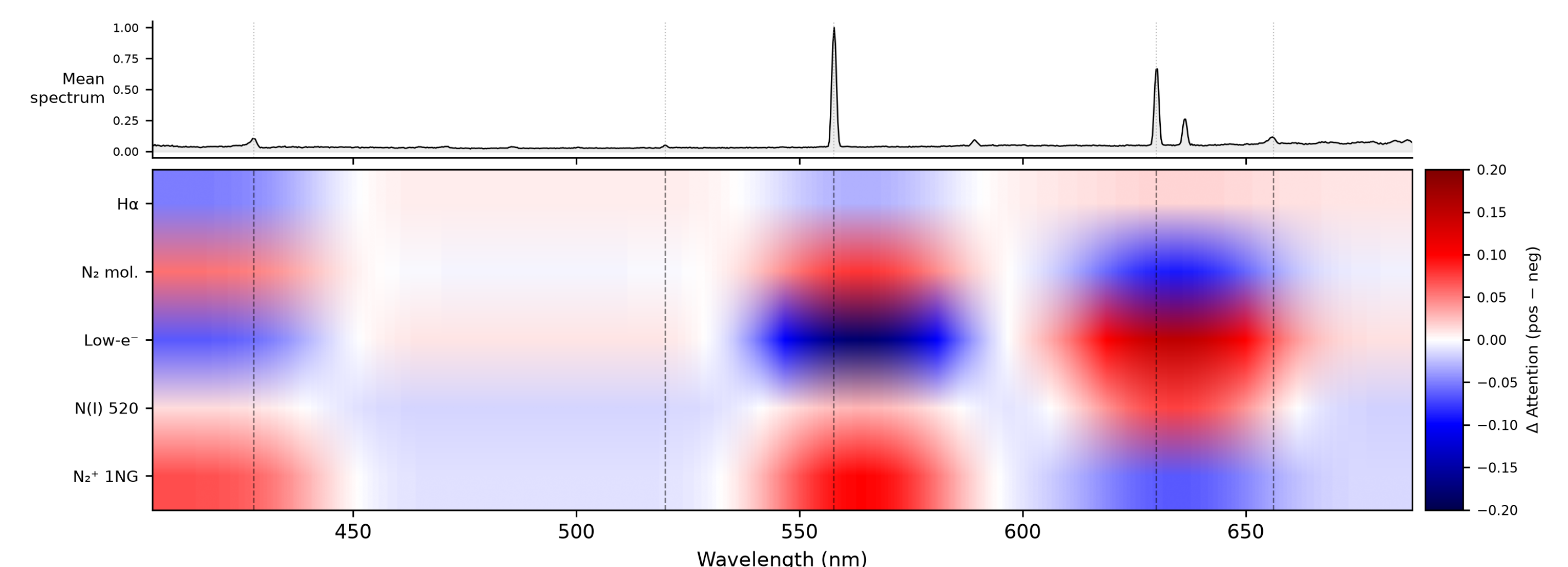


Figure 4. Differential attention $\Delta A_c(\lambda)$ per class, averaged over the 15 cross-validation models. Red: attention *increased* on positive samples; blue: *decreased*. Dashed lines mark known emission wavelengths; top: mean spectrum.

For the trained ViT-1D, the maps recover **known physical patterns** from *classification supervision alone*:

- ✓ N₂⁺ 1NG peaks at **427.8 nm**; H α rises toward **656 nm**; Low- e^- is broad **above 630 nm** (red OI).
- ✓ N₂ mol. is read from **co-occurring** emissions (427.8/557.7 nm), consistent with co-production by **energetic particle impact on N₂**.
- ✓ The *unlabelled* 557.7 nm line is attended by N₂ mol. and N₂⁺ 1NG, **co-excitation recovered without supervision**.
- ? N(I) 520: read via co-occurrence (OI 630/636 nm), never its own line, **needs further study**.

Contributions & Conclusion

- New 1D task**: multi-label classification on 719 expert-annotated auroral spectra (no baseline).
- Patch attention **transfers from images to 1D signals**; ViT-1D matches the MLP.
- Attention = **built-in attribution**: recovers emission lines **without priors**.
- Validated by physical ground truth, a path to **self-supervised** spectral analysis.

References

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